

Problem Set 1

1. This problem is very simple, as the  $N$  harmonic oscillators are non-interacting. So, if you know how to solve the equations of motion for one harmonic oscillator, then you have 'em all! [I.E. a many 1-body problem is easier than 1 many-body problem]. We can write the Lagrangian as:

$$L(\underline{q}, \underline{\dot{q}}, t) = \frac{1}{2} \sum_i m_i \dot{q}_i^2 - \frac{1}{2} \sum_i k_i q_i^2$$

We have the following  $N$ , Lagrange equations, 1 for each  $i$ :

$$\frac{\partial}{\partial q_i} L - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) = 0 \quad ; \quad \forall i$$

Do the differentiation and we have:

$$-k_i q_i - m_i \ddot{q}_i = 0$$

$$\text{or, } \ddot{q}_i + \omega_i^2 q_i = 0, \quad \text{where } \omega_i^2 = \frac{k_i}{m_i}$$

This is a standard differential equation, and we obtain:

$$q_i(t) = A_i \cos(\omega_i t) + B_i \sin(\omega_i t) \quad ; \quad \forall i = 1 \dots N$$

We are also given initial conditions of:

$$q_i(0) = x_i = A_i$$

$$\dot{q}_i(0) = v_i = \omega_i B_i$$

$$q_i(t) = x_i \cos(\omega_i t) + \left( \frac{v_i}{\omega_i} \right) \sin(\omega_i t) \quad ; \quad \forall i = 1 \dots N$$

The above is the equation of motion for the  $i^{\text{th}}$  oscillator and is the same for the other oscillators

2. The Lagrangian is:

$$L = a\dot{q}_1^2 + b\frac{\dot{q}_2}{q_1} + c\dot{q}_1\dot{q}_2 + f q_1^2 \dot{q}_1 q_3 + g\dot{q}_2 - k\sqrt{q_1^2 + q_2^2}$$

We can get the momenta:

$$p_1 = \frac{\partial}{\partial \dot{q}_1} L = 2a\dot{q}_1 + c\dot{q}_2 + f q_1^2 q_3 \quad (1)$$

$$p_2 = \frac{\partial}{\partial \dot{q}_2} L = \frac{b}{q_1} + c\dot{q}_1 + g \quad (2)$$

$$p_3 = \frac{\partial}{\partial \dot{q}_3} L = 0$$

The Hamiltonian is given by:

$$\mathcal{H} = p_1 \dot{q}_1 + p_2 \dot{q}_2 - L$$

$$= 2a\dot{q}_1^2 + c\dot{q}_1\dot{q}_2 + f q_1^2 \dot{q}_1 q_3 + \frac{b}{q_1} \dot{q}_2 + c\dot{q}_1\dot{q}_2 + g\dot{q}_2 - L$$

$$\mathcal{H} = a\dot{q}_1^2 + c\dot{q}_1\dot{q}_2 + k\sqrt{q_1^2 + q_2^2}$$

Eliminating  $\dot{q}_1$  and  $\dot{q}_2$  from relations (1) and (2) above:

$$\mathcal{H} = \frac{-g}{c^2} \left( p_2 - \frac{b}{q_1} - g \right)^2 + \left( p_2 - \frac{b}{q_1} - g \right) \left( p_1 - f q_1^2 q_3 \right) + k\sqrt{q_1^2 + q_2^2}$$

3. The 1-D oscillator (we've seen much of this in class):

a) 
$$H = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 q^2$$

$$H = \frac{1}{2m} (p^2 + m^2 \omega^2 q^2)$$

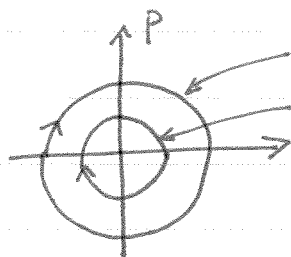
this describes an oval in  $p$ - $q$  space, but if we scale our  $q$  coordinate:

$$q' = m \omega q$$

Then we have:

$$H = \frac{1}{2m} (p^2 + (q')^2)$$

← and this is a circle



energy is fixed on this circle  
this circle is an oscillator with a different total energy  $E$ .

b) Hamilton's equations are  $\dot{q}_i = \frac{\partial H}{\partial p_i}$  and  $\dot{p}_i = -\frac{\partial H}{\partial q_i}$

For this system:

$$\frac{\partial H}{\partial p_i} = \frac{p}{m}$$

$$\frac{\partial H}{\partial q_i} = m \omega^2 q$$

Therefore:

$$\dot{q}_i = \dot{q} = \frac{p}{m}$$

$$\dot{p}_i = -m \omega^2 q$$

$$\boxed{\begin{aligned} \dot{q} &= \frac{p}{m} \\ \dot{p} &= -m \omega^2 q \end{aligned}}$$

c) To do this part, we must solve:

$$\dot{q} = \frac{p}{m} \quad \text{and} \quad \dot{p} = -m\omega^2 q$$

there's a couple of ways of doing this

1) since  $\dot{q} = p/m$   $\left. \begin{array}{l} \text{take } \text{time} \\ \text{derivative} \end{array} \right\}$   
 $\ddot{q} = \dot{p}/m$   $\left. \begin{array}{l} \text{plug in } \dot{p} \text{ formula} \end{array} \right\}$

$$\ddot{q} = -m\omega^2 q / m = -\omega^2 q$$

$\ddot{q} + \omega^2 q = 0$   $\left. \begin{array}{l} \text{solve by plugging} \\ \text{in a known} \\ \text{solution} \end{array} \right\}$

$$q(t) = A \cos \omega t + B \sin \omega t$$

$$\dot{q}(t) = -\omega A \sin \omega t + \omega B \cos \omega t$$

$$\begin{aligned} \ddot{q}(t) &= -\omega^2 A \cos \omega t - \omega^2 B \sin \omega t \\ &= -\omega^2 q(t) \end{aligned}$$

this solution works

Alternatively:

2) In matrix form:

$$\begin{pmatrix} dq/dt \\ dp/dt \end{pmatrix} = \begin{pmatrix} 0 & 1/m \\ -m\omega^2 & 0 \end{pmatrix} \begin{pmatrix} q \\ p \end{pmatrix}$$

or:

$$\underline{\dot{X}} = \underline{A} \underline{X}$$

In general, we solve these matrix equations by diagonalizing

$$\underline{U} \underline{A} \underline{U}^{-1} = \underline{\Lambda}$$

← a diagonal matrix containing eigenvalues of A

$\underline{U}$  is a unitary matrix where the column vectors are eigenvectors of  $\underline{A}$

In general; we want to define new variables

$$\underline{x}' = \underline{U} \cdot \underline{x}$$

If we take our diff eq. expression and multiply by  $\underline{U}$ , we get

$$\underline{U} \cdot \underline{\dot{x}} = \underline{U} \underline{A} \cdot \underline{x}$$

$$\underline{\dot{x}'} = \underline{U} \underline{A} \cdot \underline{x}$$

$$\underline{\dot{x}'} = \underline{U} \underline{A} \underbrace{\underline{U}^{-1} \underline{U}}_{=I} \underline{x}$$

unitary matrices  
can be inserted here

$$\underline{\dot{x}'} = \underline{\lambda} \underline{x}'$$

this is now a  
diagonal matrix  
 $\begin{pmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{pmatrix}$

which means we have uncoupled 1<sup>st</sup> order equations

$$\dot{x}'_a = \lambda_a x'_a$$

$$\frac{dx'_a}{dt} = \lambda_a x'_a$$

$$\frac{dx'_a}{x'_a} = \lambda_a dt$$

$$\ln x'_a \stackrel{\text{integrate}}{\Downarrow} = \lambda_a t \implies x'_a = e^{\lambda_a t} x'_a(0)$$

That's the most general way to solve coupled 1<sup>st</sup> order Diff Eq's.

Now, to our problem:

$$A = \begin{pmatrix} 0 & \frac{1}{m} \\ -m\omega^2 & 0 \end{pmatrix} \quad \underline{\lambda} = \begin{pmatrix} -i\omega \\ +i\omega \end{pmatrix}$$

$$\underline{U} = \begin{pmatrix} i/m\omega & 1 \\ -i/m\omega & 1 \end{pmatrix}$$

time dependent solutions arising from eigenvalues  $\underline{\lambda}$ :

$$\underline{D}(t) = \begin{pmatrix} e^{-i\omega t} & 0 \\ 0 & e^{+i\omega t} \end{pmatrix}$$

$$x'_\alpha(t) = e^{-i\omega t} x'_\alpha(0)$$

$$x'_\beta(t) = e^{+i\omega t} x'_\beta(0)$$

These say the same thing!

We can back-solve to find  $q(t)$  &  $p(t)$

$$\begin{pmatrix} q(t) \\ p(t) \end{pmatrix} = \begin{pmatrix} \frac{-im\omega}{2} & \frac{im\omega}{2} \\ \frac{1}{2} & \frac{1}{2} \end{pmatrix} \begin{pmatrix} e^{-i\omega t} & 0 \\ 0 & e^{+i\omega t} \end{pmatrix} \begin{pmatrix} \frac{i}{m\omega} & 1 \\ \frac{-i}{m\omega} & 1 \end{pmatrix} \begin{pmatrix} q(0) \\ p(0) \end{pmatrix}$$

$$\underline{X}(t) = \underline{U}^{-1} \underline{D}(t) \underline{U} \underline{X}(0)$$

transformation to  $\begin{pmatrix} q \\ p \end{pmatrix}$  space

$\frac{1}{2}$  solutions in eigenvalue space

$$\begin{pmatrix} q(t) \\ p(t) \end{pmatrix} = \begin{pmatrix} \cos \omega t & \frac{1}{m\omega} \sin \omega t \\ -m\omega \sin \omega t & \cos \omega t \end{pmatrix} \begin{pmatrix} q(0) \\ p(0) \end{pmatrix}$$

Therefore:

$$\begin{aligned} q(t) &= q_0 \cos \omega t + \frac{P_0}{m\omega} \sin \omega t \\ p(t) &= -m\omega q_0 \sin \omega t + P_0 \cos \omega t \end{aligned}$$

$$d) \quad H(P_0, q_0) = \frac{P_0^2}{2m} + \frac{1}{2} m \omega^2 q_0^2$$

$$H(p(t), q(t)) = \frac{(-m\omega q_0 \sin \omega t + P_0 \cos \omega t)^2}{2m} + \frac{1}{2} m \omega^2 \left( q_0 \cos \omega t + \frac{P_0}{m\omega} \sin \omega t \right)^2$$

$$= \frac{1}{2m} \left[ m^2 \omega^2 q_0^2 \sin^2 \omega t + P_0^2 \cos^2 \omega t - 2m\omega q_0 P_0 \sin \omega t \cos \omega t \right]$$

$$+ \frac{m\omega^2}{2} \left[ q_0^2 \cos^2 \omega t + \frac{P_0^2}{m^2 \omega^2} \sin^2 \omega t + 2 \frac{q_0 P_0}{m\omega} \cos \omega t \sin \omega t \right]$$

$$= \frac{P_0^2}{2m} \left[ \cos^2 \omega t + \sin^2 \omega t \right] + \frac{1}{2} m \omega^2 q_0^2 \left[ \sin^2 \omega t + \cos^2 \omega t \right]$$

$$+ q_0 P_0 \omega \left[ \cos \omega t \sin \omega t - \sin \omega t \cos \omega t \right]$$

$$= \frac{P_0^2}{2m} + \frac{1}{2} m \omega^2 q_0^2 \quad \checkmark$$

$$= H(P_0, q_0)$$



3 e.)

$$p = \sqrt{2m\omega J} \cos\theta$$

$$\rightarrow p^2 = 2m\omega J \cos^2\theta$$

$$q = \sqrt{\frac{2J}{m\omega}} \sin\theta$$

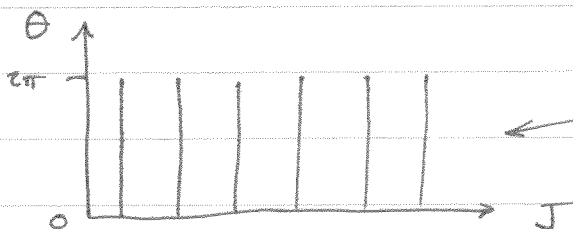
$$q^2 = \frac{2J}{m\omega} \sin^2\theta$$

$$H = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 q^2$$

$$= \omega J \cos^2\theta + \omega J \sin^2\theta = \omega J [\cos^2\theta + \sin^2\theta]$$

$$\boxed{H = \omega J}$$

f.)



phase space is a series of lines depending on initial conditions.

g.)

$$\frac{\partial H}{\partial J} = \dot{\theta}$$

$$\frac{\partial H}{\partial \theta} = -\dot{J}$$

$$\therefore \dot{\theta} = \omega$$

$$\dot{J} = 0$$

$$\therefore \boxed{\theta(t) = \theta(0) + \omega t \quad J(t) = J(0)}$$

Each line in the  $p-q$  phase space maps to a single line in the  $J-\theta$  phase space.

4. a)

$$T = \frac{1}{2m} p_1^2 + \frac{1}{2m} p_2^2 + \frac{1}{2m} p_3^2$$

(also

$$T = \frac{1}{2} m \dot{x}_1^2 + \frac{1}{2} m \dot{x}_2^2 + \frac{1}{2} m \dot{x}_3^2)$$

b)

$$V(x_1, x_2, x_3) = \frac{1}{2} k (x_2 - x_1 - r_0)^2 + \frac{1}{2} k (x_3 - x_2 - r_0)^2$$

↑

1-2 bond

↑

2-3 bond

c)

If

$$u = x_2 - x_1 - r_0$$

⇒

$$\dot{u} = \dot{x}_2 - \dot{x}_1$$

$$v = x_3 - x_2 - r_0$$

⇒

$$\dot{v} = \dot{x}_3 - \dot{x}_2$$

$$w = \frac{1}{3} (x_1 + x_2 + x_3)$$

⇒

$$\dot{w} = \frac{1}{3} (\dot{x}_1 + \dot{x}_2 + \dot{x}_3)$$

$$V(u, v, w) = \frac{1}{2} k u^2 + \frac{1}{2} k v^2 = \frac{1}{2} k (u^2 + v^2)$$

T is harder:

$$\dot{u}^2 = \dot{x}_2^2 + \dot{x}_1^2 - 2\dot{x}_1\dot{x}_2$$

$$\dot{v}^2 = \dot{x}_3^2 + \dot{x}_2^2 - 2\dot{x}_3\dot{x}_2$$

}

adding these  
won't give you T!

Write it as a matrix equation.

$$\begin{pmatrix} \dot{u} \\ \dot{v} \\ \dot{w} \end{pmatrix} = \begin{pmatrix} -1 & 1 & 0 \\ 0 & -1 & 1 \\ \frac{1}{3} & \frac{1}{3} & \frac{1}{3} \end{pmatrix} \begin{pmatrix} \dot{x}_1 \\ \dot{x}_2 \\ \dot{x}_3 \end{pmatrix}$$

To solve for  $\dot{x}_1, \dot{x}_2, \dot{x}_3$  in terms of  $\dot{u}, \dot{v}, \dot{w}$ :

$$\vec{y} = \underline{A} \cdot \vec{x}$$

$$\underline{A}^{-1} \vec{y} = \underline{A}^{-1} \underline{A} \cdot \vec{x}$$

$$\underline{A}^{-1} \vec{y} = \vec{x}$$

This gives us:

$$\dot{x}_1 = \dot{w} - \frac{2}{3}\dot{u} - \frac{1}{3}\dot{v}$$

$$\dot{x}_2 = \dot{w} + \frac{1}{3}\dot{u} - \frac{1}{3}\dot{v}$$

$$\dot{x}_3 = \dot{w} + \frac{1}{3}\dot{u} + \frac{2}{3}\dot{v}$$

Multiplying  $\dot{x}_1^2 + \dot{x}_2^2 + \dot{x}_3^2$  out we get:

$$T(u, v, w) = \frac{1}{6m} (2\dot{u}^2 + 2\dot{u}\dot{v} + 2\dot{v}^2 + 9\dot{w}^2)$$

$$V(u, v, w) = \frac{1}{2}k(u^2 + v^2)$$

$$\mathcal{H} = \frac{1}{6m} (2\dot{u}^2 + 2\dot{u}\dot{v} + 2\dot{v}^2 + 9\dot{w}^2) + \frac{1}{2}k(u^2 + v^2)$$

(Note that there is no  $w$  in the potential, so  $w$  is a ~~free~~<sup>conserved</sup> coordinate, and  $w$  is linearly growing in time.)

There's still the pesky  $uv$  coupling term in the kinetic energy.

This is a local mode picture of molecular vibrations.

We can also use normal modes:

4d)

To obtain the normal coordinates, one uses the rows of the Eigenvector matrix  $\underline{u}$ :

translation:

$$\omega_1 = 0$$

$$\vec{u}_1 = x_1 + x_2 + x_3$$

zero frequency: not vibration

$$\begin{array}{ccc} \rightarrow & \rightarrow & \rightarrow \\ 0 & m & m & 0 & m & 0 \end{array}$$

Symmetric stretch:

$$\omega_2 = \omega$$

$$\vec{u}_2 = -x_1 + x_3$$

$$\begin{array}{ccc} \leftarrow & & \rightarrow \\ 0 & m & 0 & m & 0 \end{array}$$

asymmetric stretch:

$$\omega_3 = \sqrt{3}\omega$$

$$\vec{u}_3 = x_1 - 2x_2 + x_3$$

$$\begin{array}{ccc} \rightarrow & \leftarrow & \rightarrow \\ 0 & m & 0 & m & 0 \end{array}$$

We can also reverse this to obtain the Cartesian coords in terms of the normal modes

$$x_1 = \frac{1}{6}(2u_1 - 3u_2 + u_3)$$

$$x_2 = \frac{1}{3}(u_1 - u_3)$$

$$x_3 = \frac{1}{6}(2u_1 + 3u_2 + u_3)$$

$$V = \frac{1}{2}k(x_2 - x_1 - r_0)^2 + \frac{1}{2}k(x_3 - x_2 - r_0)^2$$

$$V = \frac{k}{4}(4r_0^2 - 4r_0u_2 + u_2^2 + u_3^2) \quad \leftarrow \text{uncoupled in } V$$

$$T = \frac{m}{12}(2\dot{u}_1^2 + 3\dot{u}_2^2 + \dot{u}_3^2) \quad \leftarrow \text{uncoupled in } T$$

4d) Normal Modes: (i.e. symmetric stretch & asymmetric stretch)

To find these we must find a set of coordinates in which the force constant matrix is diagonal:

$$\underline{\omega}^2 = \underline{U} \underline{K} \underline{U}^{-1}$$

$\swarrow$  diagonal matrix of frequencies  $\nwarrow$

$$\underline{K} = \begin{pmatrix} \frac{\partial^2 V}{\partial x_1^2} \frac{1}{m_1} & \frac{\partial^2 V}{\partial x_1 \partial x_2} \frac{1}{\sqrt{m_1 m_2}} & \\ \frac{\partial^2 V}{\partial x_2 \partial x_1} \frac{1}{\sqrt{m_2 m_1}} & \frac{\partial^2 V}{\partial x_2^2} \frac{1}{m_2} & \\ & & \dots \end{pmatrix}$$

$\underline{K}$  is called a Hessian matrix or a mass-weighted force constant matrix.

For our system:

$$\underline{K} = \frac{1}{m} \begin{pmatrix} k & -k & 0 \\ -k & 2k & -k \\ 0 & -k & k \end{pmatrix}$$

$$= \frac{k}{m} \begin{pmatrix} 1 & -1 & 0 \\ -1 & 2 & -1 \\ 0 & -1 & 1 \end{pmatrix} = \omega^2 \begin{pmatrix} 1 & -1 & 0 \\ -1 & 2 & -1 \\ 0 & -1 & 1 \end{pmatrix}$$

So we need the eigenvalues ( $\underline{\omega}^2$ ) and eigenvectors ( $\underline{U}$ ) of this matrix

$$\underline{\omega}^2 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \omega^2 & 0 \\ 0 & 0 & 3\omega^2 \end{pmatrix} \quad \underline{U} = \begin{pmatrix} 1 & 1 & 1 \\ -1 & 0 & 1 \\ 1 & -2 & 1 \end{pmatrix}$$

5. This is a famous unsolved problem. I just wanted to see who could tackle it and take it the furthest:

$$H = \frac{1}{2}(p_1^2 + p_2^2 + q_1^2 + q_2^2) + q_1 q_2^2 - \frac{1}{3} q_1^3$$

$$p_1 = f_1(J_1) \cos \theta_1$$

$$q_1 = f_1(J_1) \sin \theta_1$$

$$p_2 = f_2(J_2) \cos \theta_2$$

$$q_2 = f_2(J_2) \sin \theta_2$$

$$H = \frac{1}{2}(f_1(J_1)^2 + f_2(J_2)^2) + f_1(J_1) f_2(J_2)^2 \sin \theta_1 \sin^2 \theta_2 - \frac{1}{3} f_1(J_1)^3 \sin^3 \theta_1$$

$$p_i = q_i \cot \theta_i = \frac{\partial G_i}{\partial q_i}$$

$$G_i = \frac{1}{2} q_i^2 \cot \theta_i$$

$$J_i = -\frac{\partial G_i}{\partial \theta_i} = \frac{q_i^2}{2 \sin^2 \theta_i}$$

Thus:  $q_i = \sqrt{2J_i} \sin \theta_i$

So:

$$f_i(J_i) = \sqrt{2J_i}$$

Hence:

$$H = \frac{1}{2}(2J_1 + 2J_2) + 2\sqrt{2J_1} J_2 \sin \theta_1 \sin^2 \theta_2 - \frac{1}{3} \sqrt{2J_1}^3 \sin^3 \theta_1$$

$$= J_1 + J_2 + 2\sqrt{2} J_1^{1/2} J_2 \sin \theta_1 \sin^2 \theta_2 - \frac{2\sqrt{2}}{3} J_1^{3/2} \sin^3 \theta_1$$

$$H_0 = J_1 + J_2$$

$$V = 2\sqrt{2} J_1^{1/2} J_2 \sin \theta_1 \sin^2 \theta_2 - \frac{2\sqrt{2}}{3} J_1^{3/2} \sin^3 \theta_1$$

With no perturbation, there would be two frequencies  $\omega_1$  and  $\omega_2$  such that

$$\omega_1 = \frac{\partial \mathcal{H}_0}{\partial J_1} \quad \omega_2 = \frac{\partial \mathcal{H}_0}{\partial J_2}$$

Since  $V$  is periodic in  $\theta_1$  and  $\theta_2$  we can rewrite  $gV$  as a Fourier series:

$$gV = \sum_{n_1, n_2} gV_{n_1, n_2}(J_1, J_2) e^{i(n_1 \theta_1 + n_2 \theta_2)}$$

Moser's 1974 paper [Moser, J in The Stability of the Solar System (ed. Y. Kozai) pp. 1-9, International Astronomical Union] has a long discussion about perturbation theories and where to go next. . . .